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# Two-loop self-energy corrections to the fine structure 

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Received 18 October 2001, in final form 31 December 2001
Published 15 February 2002
Online at stacks.iop.org/JPhysA/35/1927


#### Abstract

We investigate two-loop higher order binding corrections to the fine structure, which contribute to the spin-dependent part of the Lamb shift. Our calculation focuses on the so-called 'two-loop self-energy' involving two virtual closed photon loops. For bound states, this correction has proved to be notoriously difficult to evaluate. The calculation of the binding corrections to the boundstate two-loop self-energy is simplified by a separate treatment of hard and soft virtual photons. The two photon-energy scales are matched at the end of the calculation. We explain the significance of the mathematical methods employed in the calculation in a more general context, and present results for the fine-structure difference of the two-loop self-energy through the order of $\alpha^{8}$.


PACS numbers: 31.15.-p, 12.20.Ds, 31.30Jv, 32.10.Fn

## 1. Introduction

Ultra-precise measurements in atomic systems represent today one of the most stringent available tests of fundamental quantum theories and a means for the determination of fundamental physical constants with unprecedented accuracy [1]. The theoretical description of the bound states at a level of accuracy which matches the current experimental precision, which has reached 1.8 parts in $10^{14}$ and whose accuracy is to be improved in the near future [2], demands a thorough understanding of the bound state including-among other effectsthe relativistic, one-loop, two-loop and higher-order radiative, recoil, radiative-recoil and nuclear-size corrections [3, 4].

We focus here on radiative corrections, which can be described-for atomic systems with low nuclear charge number-by a nonanalytic expansion in powers of the three parameters (i) $\alpha$ (the fine-structure constant), (ii) the product $Z \alpha$ ( $Z$ is the nuclear charge number) and (iii) the logarithm $\ln \left[(Z \alpha)^{-2}\right]$. The expansion in powers of $\alpha$, which is


Figure 1. Feynman diagrams representing the two-photon electron self-energy. The double line denotes the bound electron propagator. The arrow of time is from right to left.
the perturbation theory parameter in quantum electrodynamics (QED), corresponds to the number of loops in the diagrams. The bound-state effects are taken into account by the expansions in the two latter parameters. Higher order terms in the expansions in powers of $Z \alpha$ and $\ln \left[(Z \alpha)^{-2}\right]$ are referred to as the 'binding corrections'. One of the historically most problematic sets of Feynman diagrams in the treatment of the Lamb shift for atomic systems has been the radiative correction due to two closed virtual-photon loops shown in figure 1.

Let us recall at this point that even the evaluation of higher order binding corrections to the one-loop self-energy, which a priori should represent a less involved calculational challenge, has represented a problem for analytic evaluations for over three decades [5-9]. The energy shifts of the bound states due to the radiative corrections are conveniently expressed by expansion coefficients corresponding to the powers of $Z \alpha$ and $\ln \left[(Z \alpha)^{-2}\right]$; the naming convention is that the power of $Z \alpha$ and the power of the logarithm are indicated as indices to the analytic coefficients (see also equation (1)) below. Because the expansion in both the one-loop and two-loop cases starts with the fourth power of $Z \alpha$, the non-vanishing coefficients carry indices $A_{k l}$ and $B_{k l}$ for the one- and two-loop cases, respectively (with $k \geqslant 4$-see [3] for a comprehensive review).

Logarithmic corrections with $l \geqslant 1$ can sometimes be inferred separately in a much simplified approach, e.g. by considering infrared divergent contributions to electron form factors. By contrast, the higher order non-logarithmic coefficients represent a considerable calculational challenge. Realistically, i.e. with the help of current computer algebra systems $[10]^{4}$, one can hope to evaluate non-logarithmic coefficients of sixth order in $Z \alpha$. Complete results for the one-loop higher order correction $A_{60}$ for S and P states have only been available recently $[8,11,12]$. Calculational difficulties have until now precluded a successful evaluation of the corresponding coefficient $B_{60}$ for the two-loop effect. The groundwork for the evaluation of $B_{60}$ was laid in [13]. Here, we are concerned with the evaluation of the fine-structure differences of the logarithmic and non-logarithmic coefficients $B_{6 L}$ (where $L=0,1,2$ ), i.e. with the $n \mathrm{P}_{3 / 2}-n \mathrm{P}_{1 / 2}$ difference of these coefficients.

[^0]Using natural Gaussian units ( $\hbar=c=\epsilon_{0}=1$ ), as is customary for the current type of calculation, we write the two-photon self-energy in the $Z \alpha$-expansion for P states in terms of $B$-coefficients as
$\Delta E_{\mathrm{SE}}=\left(\frac{\alpha}{\pi}\right)^{2}(Z \alpha)^{4} \frac{m}{n^{3}}\left[B_{40}+(Z \alpha)^{2}\left[B_{62} \ln ^{2}(Z \alpha)^{-2}+B_{61} \ln (Z \alpha)^{-2}+B_{60}\right]+\mathcal{R}\right]$
where the remainder $\mathcal{R}$ is of order $o(Z \alpha)^{3}$. Relevant Feynman diagrams are shown in figure 1.
Here, $m$ denotes the electron mass (we write equation (1) in the non-recoil limit, i.e. for an infinite nuclear mass). The double logarithmic $B_{62}$-coefficient is spin-independent, so that we have $\Delta_{\mathrm{fs}} B_{62}=0$. In this paper, we evaluate the fine-structure differences
$\Delta_{\mathrm{fs}} B_{61}=B_{61}\left(n \mathrm{P}_{3 / 2}\right)-B_{61}\left(n \mathrm{P}_{1 / 2}\right) \quad \Delta_{\mathrm{fs}} B_{60}=B_{60}\left(n \mathrm{P}_{3 / 2}\right)-B_{60}\left(n \mathrm{P}_{1 / 2}\right)$.
Throughout the paper, we will follow the convention that $\Delta_{\mathrm{fs}} X \equiv X\left(n \mathrm{P}_{3 / 2}\right)-X\left(n \mathrm{P}_{1 / 2}\right)$ denotes the 'fine-structure part' of a given quantity $X$. For $\Delta_{\mathrm{fs}} B_{61}$ and $\Delta_{\mathrm{fs}} B_{60}$, we provide complete results. It is perhaps worth noting that two-loop self-energy effects for bound states have represented a considerable challenge for theoretical evaluations. Our investigation represents a continuation of the previous study on the two-loop problem (see e.g. [13-16]). It is probably a triviality to express that technical difficulties in the calculation and its description in the following sections of the paper cannot be avoided.

For the description of the self-energy radiative effects, mediated by hard virtual photons, we use the modified Dirac Hamiltonian
$H_{\mathrm{D}}^{(m)}=\boldsymbol{\alpha} \cdot\left[\boldsymbol{p}-\mathrm{e} F_{1}(\Delta) \boldsymbol{A}\right]+\beta m+\mathrm{e} F_{1}(\Delta) \phi+F_{2}(\Delta) \frac{e}{2 m}(\mathrm{i} \gamma \cdot \boldsymbol{E}-\beta \boldsymbol{\sigma} \cdot \boldsymbol{B})$
which approximately describes an electron subject to an external scalar potential $\phi \equiv \phi(\boldsymbol{r})$ and an external vector potential $\boldsymbol{A} \equiv \boldsymbol{A}(\boldsymbol{r})$. This modified Hamiltonian is still local in coordinate space. The Dirac matrices in (3) are to be understood in the standard (Dirac) representation [17] (in the following, we will also use the non-covariant notation $\beta \equiv \gamma^{0}$ and $\alpha^{i} \equiv \gamma^{0} \gamma^{i}$ ).

The argument $\Delta$ of the electron form factors $F_{1}$ and $F_{2}$ in equation (3) is to be interpreted as a Laplacian operator acting on all quantities to the right (but not on the wavefunction of the bound electron in evaluating $\left.H_{\mathrm{D}}^{(m)}|\psi\rangle\right)$. In momentum space, the action of the Hamiltonian $H_{\mathrm{D}}^{(m)}$ is described by the convolution $\left[H_{\mathrm{D}}^{(m)} \psi\right]\left(\boldsymbol{p}^{\prime}\right)=\int \mathrm{d}^{3} p /(2 \pi)^{3} H_{\mathrm{D}}^{(m)}\left(\boldsymbol{p}^{\prime}-\boldsymbol{p}\right) \psi(\boldsymbol{p})$. The form factors-in momentum space-assume arguments according to the replacement $\Delta \rightarrow-q^{2} \equiv-\left(p^{\prime}-p\right)^{2}$. In equation (3), radiative corrections are taken into account in the sense of an effective theory via the inclusion of the on-shell form factors $F_{1}$ and $F_{2}$. Although the bound electron is not an on-shell particle, the modified Hamiltonian (3) can still approximately account for significant radiative systems with low nuclear charge number $Z$. Of course, the Hamiltonian (3) cannot offer a complete description of the bound electron. Recoil effects cannot be described by a one-particle equation in principle, and vacuum-polarization effects are not contained in equation (3). However, the effective description of self-energy radiative corrections mediated by hard virtual photons given by equation (3) will turn out to be useful in the context of the current investigation.

Both the form factors $F_{1}$ and $F_{2}$ entering in equation (3) are infrared divergent, but this divergence is cut off in a natural way at the atomic binding energy scale $(Z \alpha)^{2} m$. The fact that on-shell form factors can describe radiative corrections to the fine structure-mediated by high-energy virtual photons-has been demonstrated explicitly in [18]. The modified Dirac Hamiltonian (3) and the associated modified Dirac equation have been introduced-in the one-loop approximation-in chapter 7 of [17] (see e.g. equations (77) and (103) of [17] (chapter 7)). The low-energy part of the calculation is carried out using nonrelativistic approximations in the spirit of the simplified treatment introduced in the previous oneand two-loop calculations [8, 11-13, 19]. This approach was inspired, in part, by various
attempts to formulate simplified low-energy (nonrelativistic) approximations to quantum electrodynamics ('NRQED'), see e.g. [20, 21]. Both the high-energy and the low-energy contributions are matched at the separation scale $\epsilon$ whose role in the calculation is illustrated by the mathematical model example discussed in appendix A.

In a two-loop calculation, either of the two virtual photons may have a high or low energy as compared to the separation scale $\epsilon$. A priori, this necessitates [13] a separation of the calculation into three different contributions: (i) both photon energies large, (ii) one photon with a large and one with a small energy and (iii) both photons with small energies. For the particular problem at hand (the fine-structure differences of $B_{61}$ and $B_{60}$ ), we are in the fortunate position that effects caused by hard virtual photons (i) are described by the modified Dirac Hamiltonian (3), whereas the low-energy part discussed in section 4 below comprises both remaining contributions (ii) and (iii).

This paper is organized as follows: two-loop form factors entering in equation (3) are analysed in section 2. The calculation is split into two parts: the high-energy part discussed in section 3 and the low-energy part, which is treated along ideas introduced in [21] in section 4. Results and conclusions are left to section 5.

## 2. Two-loop form factors

In order to analyse the modified Dirac Hamiltonian (3) through two-loop order, we first have to investigate certain expansion coefficients of the electronic $F_{1}$ and $F_{2}$ form factors which are thoroughly discussed in the seminal papers [22,23]. For the momentum transfer $q^{2}$ which is the argument of the two functions $F_{1} \equiv F_{1}\left(q^{2}\right)$ and $F_{2} \equiv F_{2}\left(q^{2}\right)$, we use the convention $q^{2}=q_{\mu} q^{\mu}=\left(q^{0}\right)^{2}-\boldsymbol{q}^{2}$. The variable $t$ in [22,23] is given as $t=q^{2}$. When we evaluate radiative corrections to the binding Coulomb field which is mediated by spacelike virtual photons, we have $q^{2}=-\boldsymbol{q}^{2}$ because $q^{0}=0$. We use the conventions (see equation (1.2) in [22]):

$$
\begin{equation*}
F_{1}(t)=1+\sum_{n=1}^{\infty}\left(\frac{\alpha}{\pi}\right)^{n} F_{1}^{(2 n)}(t) \quad F_{2}(t)=\sum_{n=1}^{\infty}\left(\frac{\alpha}{\pi}\right)^{n} F_{2}^{(2 n)}(t) \tag{4}
\end{equation*}
$$

One- and two-loop effects are denoted by upper indices 2 and 4, respectively. This notation is motivated by the observation that two-loop effects are of fourth order in the quantum electrodynamic interaction Lagrangian $-e \bar{\psi} \gamma^{\mu} A_{\mu} \psi$ (in the Furry picture, which is used for the description of bound states, the Coulomb interaction is taken out of the interaction Lagrangian).

There are two different points of view regarding the choice of diagrams to be included in the two-loop form factors, depending on whether the self-energy vacuum polarization diagram of figure 2 is included in the calculation or not. We will discuss both cases and give results with and without the diagram shown in figure 2 taken into account.

First, we discuss the results obtained for $F_{1}$ including the combined self-energy vacuum polarization diagram. In this case, the known results for the slopes $F_{1}^{\prime}(0)$ and $F_{2}(0)$, through the two-loop order, read as follows. From equation (1.11) of [22], we have
$m^{2} F_{1}^{\prime}(0)=\frac{\alpha}{\pi}\left[-\frac{1}{3} \ln \left(\frac{\lambda}{m}\right)-\frac{1}{8}\right]+\left(\frac{\alpha}{\pi}\right)^{2}\left[-\frac{4819}{5184}-\frac{49}{72} \zeta(2)+3 \zeta(2) \ln 2-\frac{3}{4} \zeta(3)\right]$
where $\lambda$ is the fictitious photon mass and the fourth-order coefficient has the numerical value

$$
\begin{equation*}
m^{2} F_{1}^{\prime(4)}(0)=0.469941487460 . \tag{6}
\end{equation*}
$$


(V)

Figure 2. Combined self-energy vacuum-polarization diagram (denoted ' $V$ ' in the text).

According to equation (1.7) in [22], the value of $F_{2}(0)$, through two-loop order, reads

$$
\begin{equation*}
F_{2}(0)=\frac{1}{2} \frac{\alpha}{\pi}+\left(\frac{\alpha}{\pi}\right)^{2}\left[\frac{197}{144}+\frac{1}{2} \zeta(2)-3 \zeta(2) \ln 2+\frac{3}{4} \zeta(3)\right] \tag{7}
\end{equation*}
$$

where the two-loop coefficient has the numerical value

$$
\begin{equation*}
F_{2}^{(4)}(0)=-0.328478965579 . \tag{8}
\end{equation*}
$$

We now turn to the discussion of the slope $F_{2}^{\prime(4)}(0)$. In view of equation (1.20) of [22] (see also [24]), we have (up to two-loop order)

$$
\begin{equation*}
F_{2}(t)=\frac{\alpha}{\pi} \mathcal{F}_{2}^{(2)}(t)+\left(\frac{\alpha}{\pi}\right)^{2}\left[\ln \frac{\lambda}{m} B(t) \mathcal{F}_{2}^{(2)}(t)+\mathcal{F}_{2}^{(4)}(t)\right] \tag{9}
\end{equation*}
$$

where the coefficients $\mathcal{F}$ are by definition infrared safe and

$$
\begin{equation*}
\mathcal{F}_{2}^{(2)}(0)=\frac{1}{2} \quad B(t)=-\frac{t}{3 m^{2}}-\frac{t^{2}}{20 m^{2}}+\mathcal{O}\left(t^{3}\right) . \tag{10}
\end{equation*}
$$

Equations (9) and (10) uniquely determine the infrared divergent contribution to $F_{2}^{(4)}(0)$. An analytic expression for $\mathcal{F}_{2}^{(4)}(t), t$ spacelike, has recently been obtained [25] in terms of harmonic polylogarithms [26, 27]. As a byproduct, an analytic expression for the slope $\mathcal{F}_{2}^{\prime(4)}(0)$ was found.

The result reads

$$
\begin{align*}
& m^{2} F_{2}^{\prime(4)}(0)=-\frac{1}{6} \ln \left(\frac{\lambda}{m}\right)+m^{2} \mathcal{F}_{2}^{\prime(4)}(0)  \tag{11}\\
& m^{2} \mathcal{F}_{2}^{\prime(4)}(0)=\frac{1751}{2160}+\frac{13}{20} \zeta(2)-\frac{23}{10} \zeta(2) \ln 2+\frac{23}{40} \zeta(3)
\end{align*}
$$

A numerical result for $\mathcal{F}_{2}^{\prime(4)}(0)$, complementing the above analytic expression, can easily be derived in combining equations (1.20), (1.30) and (3.2) in [22], as will be explained in the following. The dispersion relation (1.30) in [22] reads
$\operatorname{Re} F_{2}(t)=-\frac{4 m^{2}}{t-4 m^{2}} F_{2}(0)+\frac{1}{\pi} \frac{t}{t-4 m^{2}} P \int_{4 m^{2}}^{\infty} \frac{\mathrm{d} t^{\prime}}{t^{\prime}-t} \frac{t^{\prime}-4 m^{2}}{t^{\prime}} \operatorname{Im} F_{2}\left(t^{\prime}\right)$
where $P$ denotes the Cauchy principal value. Equation (12) applies also if we single out the two-loop effect and differentiate at zero momentum transfer, and we obtain for the slope $F_{2}^{\prime(4)}(0)$ the relation
$m^{2} F_{2}^{\prime(4)}(0)=\frac{1}{4} F_{2}^{(4)}(0)+\frac{1}{4 \pi} P \int_{4 m^{2}}^{\infty} \mathrm{d} t^{\prime} \frac{4 m^{2}-t^{\prime}}{t^{\prime 2}} \operatorname{Im}{F_{2}^{\prime(4)}\left(t^{\prime}\right)=\frac{1}{4} F_{2}^{(4)}(0)+\mathcal{T}, ~(0)}^{(1)}$
where $F_{2}^{(4)}(0)$ is given in equation (8). The second term on the right-hand side, denoted by $\mathcal{T}$, can be evaluated using the result for $\operatorname{Im} F_{2}^{(4)}(x)$ presented in equation (3.2) in [22]; it reads
$\mathcal{T}=-\int_{0}^{1} \mathrm{~d} x \frac{(1-x)^{3}}{x(1+x)} \operatorname{Im} F_{2}^{(4)}(x)=-\frac{1}{6} \ln \left(\frac{\lambda}{m}\right)+0.030740507833(1)$.
Here, the last error is due to numerical integration, and use is made of the natural variable [22]

$$
\begin{equation*}
x=\frac{1-\sqrt{1-4 m^{2} / t}}{1+\sqrt{1-4 m^{2} / t}} \tag{15}
\end{equation*}
$$

In combining the result of equation (8) with equations (13) and (14), the result $m^{2} \mathcal{F}_{2}^{\prime(4)}(0)=$ $-0.051379233561(1)$ is obtained which is in agreement with (11).

Now we will provide results for the form factors obtained excluding the self-energy vacuum-polarization graph $V$ shown in figure 2. These results refer to the pure two-photon self-energy diagrams shown in figure 1. The two-loop self-energy diagrams independently form a gauge-invariant set. They represent a historically problematic correction, and are the main subject of our investigation. The combined self-energy vacuum-polarization diagram, according to equations (1.9) and (1.10) in [23]—taking into account the subtracted dispersion relation (1.30) of [22]-leads to the following corrections:

$$
\begin{align*}
& F_{1}^{\prime(4), V}(0)=-\frac{1099}{1296}+\frac{77}{144} \zeta(2)=0.031588972474 \\
& F_{2}^{(4), V}(0)=\frac{119}{36}-2 \zeta(2)=0.015687421859  \tag{16}\\
& F_{2}^{\prime(4), V}(0)=\frac{311}{216}-\frac{7}{8} \zeta(2)=0.000497506323 .
\end{align*}
$$

For the pure self-energy graphs, which we would like to denote by the symbol $S$, we therefore obtain the following results:

$$
\begin{align*}
m^{2} F_{1}^{\prime(4), S}(0) & =-\frac{47}{576}-\frac{175}{144} \zeta(2)+3 \zeta(2) \ln 2-\frac{3}{4} \zeta(3)=0.438352514986  \tag{17}\\
F_{2}^{(4), S}(0)= & -\frac{31}{16}+\frac{5}{2} \zeta(2)-3 \zeta(2) \ln 2+\frac{3}{4} \zeta(3)=-0.344166387438  \tag{18}\\
m^{2} F_{2}^{\prime(4), S}(0) & =-\frac{1}{6} \ln \left(\frac{\lambda}{m}\right)-\frac{151}{240}+\frac{61}{40} \zeta(2)-\frac{23}{10} \zeta(2) \ln 2+\frac{23}{40} \zeta(3) \\
& =-\frac{1}{6} \ln \left(\frac{\lambda}{m}\right)-0.051876739885 \\
& \equiv-\frac{1}{6} \ln \left(\frac{\lambda}{m}\right)+\mathcal{F}_{2}^{\prime(4), S}(0) . \tag{19}
\end{align*}
$$

where the latter equality defines $\mathcal{F}_{2}^{\prime(4), S}(0)$ in analogy with equations (9) and (11).

## 3. High-energy part

Based on the modified Dirac Hamiltonian (3), corrections to the energy of the bound Dirac particle can be inferred. We will refer to the energy corrections attributable to the $F_{1}$ and $F_{2}$ form factors as $E_{1}$ and $E_{2}$, respectively. For $E_{1}$, we have

$$
\begin{equation*}
E_{1}=\left\langle\left.\left[F_{1}\left(-q^{2}\right)-1\right] e \phi\right|_{\mathrm{fs}}\right. \tag{20}
\end{equation*}
$$

where the index fs refers to the fine-structure terms, i.e. to the result obtained by subtracting the value of the matrix element for a $n \mathrm{P}_{3 / 2}$ state from the value of the same matrix element evaluated on a $n \mathrm{P}_{1 / 2}$ state. A matrix element $\langle A\rangle_{\mathrm{fs}}$ of a given operator $A$ is evaluated as

$$
\langle A\rangle_{\mathrm{fs}} \equiv\left\langle\psi_{n \mathrm{P}_{3 / 2}}^{+}\right| A\left|\psi_{n \mathrm{P}_{3 / 2}}\right\rangle-\left\langle\psi_{n \mathrm{P}_{1 / 2}}^{+}\right| A\left|\psi_{n \mathrm{P}_{1 / 2}}\right\rangle
$$

where $\psi^{+}$denotes the Hermitian conjugate of the Dirac wavefunction $\psi$ (not the Dirac adjoint $\bar{\psi}=\psi^{+} \gamma^{0}$ ). The Dirac wavefunctions $\psi$ are expanded in powers of $(Z \alpha)$ up to the order relevant for the current investigation. This expansion avoids potential problems associated with the logarithmic divergence of the Dirac wavefunction at the origin.

For $E_{1}$, up to the order of $(Z \alpha)^{6}$, we have

$$
\begin{equation*}
E_{1}=4 \pi Z \alpha F_{1}^{\prime(4)}(0)\left\langle\delta^{(3)}(\boldsymbol{r})\right\rangle_{\mathrm{fs}} \tag{21}
\end{equation*}
$$

For P states, the nonrelativistic (Schrödinger) wavefunction-the leading term in the $Z \alpha$ expansion of the Dirac wavefunction-vanishes at $r=0$, but the first relativistic correction gives a finite contribution, resulting in

$$
\begin{equation*}
\left\langle\delta^{(3)}(r)\right\rangle_{\mathrm{fs}}=-\frac{n^{2}-1}{4 n^{5}}(Z \alpha)^{5} m^{3} . \tag{22}
\end{equation*}
$$

This leads, again up to the order of $(Z \alpha)^{6}$, to the following result for $E_{1}$ :

$$
\begin{equation*}
E_{1}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{6}}{n^{3}}\left[-F_{1}^{\prime(4)}(0) \frac{n^{2}-1}{n^{2}}\right] m^{3} . \tag{23}
\end{equation*}
$$

Observe that the derivative of the $F_{1}$ form factor has a physical dimension of $1 / \mathrm{m}^{2}$ in natural units, giving the correct physical dimension for $E_{1}$. The correction due to $F_{2}$ in (3) reads

$$
\begin{equation*}
E_{2}=\left\langle F_{2}\left(-\boldsymbol{q}^{2}\right) \frac{e}{2 m} \mathrm{i} \gamma \cdot \boldsymbol{E}\right\rangle_{\mathrm{fs}} . \tag{24}
\end{equation*}
$$

A particle in an external binding Coulomb field feels an electric field $\boldsymbol{E}=\mathrm{i}(Z e) \boldsymbol{q} / \boldsymbol{q}^{2}$ —in momentum space-or $\boldsymbol{E}=-(Z e) \boldsymbol{r} /\left(4 \pi r^{3}\right)$ in coordinate space. Vacuum polarization corrections to $\boldsymbol{E}=-(Z e) \boldsymbol{r} /\left(4 \pi r^{3}\right)$ lead to higher order effects. The correction $E_{2}$ splits up in a natural way into two contributions $E_{2 a}$ and $E_{2 b}$ which are associated with $F_{2}(0)$ and the slope $F_{2}^{\prime}(0)$, respectively. $E_{2 a}$ reads

$$
\begin{equation*}
E_{2 a}=\frac{Z \alpha}{2 m} F_{2}^{(4)}(0)\left\langle-\mathrm{i} \frac{\gamma \cdot \boldsymbol{r}}{r^{3}}\right\rangle_{\mathrm{fs}} \tag{25}
\end{equation*}
$$

The evaluation of the matrix element leads to

$$
\begin{equation*}
\left\langle-\mathrm{i} \frac{\gamma \cdot \boldsymbol{r}}{r^{3}}\right\rangle_{\mathrm{fs}}=\left\{\frac{(Z \alpha)^{3}}{n^{3}}+\left[\frac{487}{360}+\frac{5}{4 n}-\frac{23}{10 n^{2}}\right] \frac{(Z \alpha)^{5}}{n^{3}}\right\} m^{2} \tag{26}
\end{equation*}
$$

For the purpose of the current investigation, the $(Z \alpha)^{6}$-component of $E_{2 a}$ is selected only:

$$
\begin{equation*}
E_{2 a}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{6}}{n^{3}}\left[F_{2}^{(4)}(0)\left(\frac{487}{720}+\frac{5}{8 n}-\frac{23}{20 n^{2}}\right)\right] m \tag{27}
\end{equation*}
$$

The matrix element $E_{2 b}$ can be expressed as

$$
\begin{equation*}
E_{2 b}=\frac{4 \pi Z \alpha}{2 m} F_{2}^{\prime(4)}(0)\langle\gamma \cdot \boldsymbol{q}\rangle_{\mathrm{fs}} . \tag{28}
\end{equation*}
$$

A transformation into coordinate space leads to

$$
\begin{equation*}
\langle\boldsymbol{\gamma} \cdot \boldsymbol{q}\rangle_{\mathrm{fs}}=\mathrm{i}\left[\frac{\partial}{\partial \boldsymbol{x}}\left(\psi^{+}(\boldsymbol{x}) \gamma \psi(\boldsymbol{x})\right)\right]_{x=0, \mathrm{fs}}=-\frac{n^{2}-1}{n^{5}}(Z \alpha)^{5} m^{4} . \tag{29}
\end{equation*}
$$

As a function of the principal quantum number $n$, the result for $E_{2 b}$ reads as follows:

$$
\begin{equation*}
E_{2 b}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{6}}{n^{3}}\left[-2 F_{2}^{\prime(4)}(0) \frac{n^{2}-1}{n^{2}}\right] m^{3} . \tag{30}
\end{equation*}
$$

This result involves the infrared divergent slope of the $F_{2}$ form factor (see equations (11) and (19)). We are thus faced with the problem of matching the infrared divergence of the slope of the $F_{2}$ form factor, expressed in terms of the fictitious photon mass $\lambda$, with the usual (energy matching parameter) $\epsilon$ introduced originally in [8]. This can be done in two ways: (i) by matching the infrared divergence of the rate of soft bremsstrahlung, calculated with a fictitious photon mass $\lambda$, to a result of the same calculation, carried out with an explicit infrared cut-off $\epsilon$ for the photon energy. This way of calculation is described in [17, pp 361-2]. It leads to the result

$$
\begin{equation*}
\ln \frac{\lambda}{2 \epsilon}=-\frac{5}{6} . \tag{31}
\end{equation*}
$$

The matching procedure (ii) consists of a comparison of the result of the application of the formalism considered above, and its application to the high-energy part of the ground state Lamb shift, which is in leading order given by the infrared divergence of the $F_{1}$ form factor, and the result obtained by direct calculation of this high-energy part in a non-covariant formalism with an explicit energy cut-off $\epsilon$, as it has been carried out in [8]. This second matching procedure leads to the following result, in agreement with (31):

$$
\begin{equation*}
\ln \frac{m}{\lambda}-\frac{3}{8}=\ln \frac{m}{2 \epsilon}+\frac{11}{24} . \tag{32}
\end{equation*}
$$

So, we are led to the replacement

$$
\begin{equation*}
-\ln \frac{\lambda}{m} \rightarrow \ln \frac{m}{2 \epsilon}+\frac{5}{6} \tag{33}
\end{equation*}
$$

A comparison with the results in equations (11), (19) and (36) reveals that the logarithmic divergence for the fine-structure difference is given by a term

$$
\begin{equation*}
-\frac{n^{2}-1}{3 n^{2}} \ln \frac{m}{2 \epsilon} \tag{34}
\end{equation*}
$$

so that we may anticipate at this stage the result for $\Delta_{\mathrm{fs}} B_{61}$

$$
\begin{equation*}
\Delta_{\mathrm{fs}} B_{61}=-\frac{n^{2}-1}{3 n^{2}} . \tag{35}
\end{equation*}
$$

Based on (30) and (33), we can express $E_{2 b}$ in terms of $\epsilon$ and $\mathcal{F}_{2}^{\prime(4)}(0)$
$E_{2 b}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{6}}{n^{3}}\left[-\frac{1}{3} \frac{n^{2}-1}{n^{2}} \ln \frac{m}{2 \epsilon}-\left(\frac{5}{18}+2 \mathcal{F}_{2}^{\prime(4)}(0) m^{2}\right) \frac{n^{2}-1}{n^{2}}\right] m$.
There is a third correction due to the effect of two one-loop corrections on the electron vertices. Because we are only interested in the fine structure, we isolate the terms which are proportional to the spin-orbit coupling, and obtain

$$
\begin{equation*}
E_{3}=\left\langle\left[2 F_{2}(0)\right] H_{\mathrm{fs}}\left(\frac{1}{E-H}\right)^{\prime}\left[2 F_{2}(0)\right] H_{\mathrm{fs}}\right\rangle_{\mathrm{fs}} \tag{37}
\end{equation*}
$$

where

$$
\begin{equation*}
H_{\mathrm{fs}}=\frac{Z \alpha}{4 m^{2} r^{3}} \sigma \cdot \boldsymbol{L} \tag{38}
\end{equation*}
$$

and $1 /(E-H)^{\prime}$ is the nonrelativistic, spin-independent reduced Schrödinger-Coulomb Green function $[28,29]$. The only spin-dependence in (37) occurs in the coupling $\sigma \cdot \boldsymbol{L}$, and it can be taken into account by an overall factor

$$
\begin{equation*}
\left\langle(\boldsymbol{\sigma} \cdot \boldsymbol{L})^{2}\right\rangle_{\mathrm{fs}}=-3 . \tag{39}
\end{equation*}
$$

We are therefore led to consider the 'spin-independent version' of the matrix element which occurs in equation (37) and obtain the following result:

$$
\begin{equation*}
\left\langle\frac{Z \alpha}{4 m^{2} r^{3}}\left(\frac{1}{E-H}\right)^{\prime} \frac{Z \alpha}{4 m^{2} r^{3}}\right\rangle_{n \mathrm{P}}=\left(-\frac{227}{8640}-\frac{1}{96 n}+\frac{1}{80 n^{2}}\right) \frac{(Z \alpha)^{6} m}{n^{3}} . \tag{40}
\end{equation*}
$$

The spin-dependence can be easily restored by considering equation (39). The index ' $n \mathrm{P}$ ' in equation (40) means that the matrix element is evaluated with the nonrelativistic, spinindependent (Schrödinger) wavefunction. Alternatively, one may evaluate with either the $n \mathrm{P}_{1 / 2}$ or the $n \mathrm{P}_{3 / 2}$ Dirac wavefunction and expand up to the leading order in $(Z \alpha)$.

The evaluation of (40) can proceed, e.g. by solving the differential equation which defines the correction to the wavefunction induced by $H_{\mathrm{fs}}$, and subsequent direct evaluation of the resulting standard integrals using computer algebra [10]. The final result for $E_{3}$ reads

$$
\begin{equation*}
E_{3}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{6}}{n^{3}}\left[\frac{227}{2880}+\frac{1}{32 n}-\frac{3}{80 n^{2}}\right] m . \tag{41}
\end{equation*}
$$

This concludes the discussion of the high-energy part. The final result for the high-energy part is

$$
\begin{equation*}
E_{\mathrm{H}}=E_{1}+E_{2 a}+E_{2 b}+E_{3} \tag{42}
\end{equation*}
$$

where $E_{1}, E_{2 a}, E_{2 b}, E_{3}$ are given in equations (23), (27), (36), (41), respectively.

## 4. Low-energy part

The low-energy part consists essentially of two contributions. Both effects, denoted here by $E_{4}$ and $E_{5}$, can be obtained by a suitable variation of the low-energy part of the oneloop self-energy, by considering the spin-dependent effects introduced by a further one-loop electron anomalous magnetic moment interaction. The first of the two terms, $E_{4}$, is caused by spin-dependent higher order effects in the one-loop self-energy, which receive additional corrections due to the anomalous magnetic moment of the electron. The second term, $E_{5}$, is due to an anomalous magnetic moment correction to the electron transition current, which can also be seen as a correction to the radiation field of the electron due to its anomalous magnetic moment.

The leading-order low-energy part (see [8]) reads

$$
\begin{equation*}
E_{\mathrm{L}}=-\frac{2 \alpha}{3 \pi m} \int_{0}^{\epsilon} \mathrm{d} \omega \omega\langle\phi| \boldsymbol{p} \frac{1}{H-(E-\omega)} \boldsymbol{p}|\phi\rangle . \tag{43}
\end{equation*}
$$

In order to isolate the fine-structure effects, we should now consider corrections to the wavefunction, current, Hamiltonian and energy of the bound state due to the spin-dependent relativistic (spin-orbit) Hamiltonian

$$
\begin{equation*}
\mathcal{H}=F_{2}(0) \frac{e}{2 m} \mathrm{i} \gamma \cdot \boldsymbol{E}=\frac{\alpha(Z \alpha)}{4 \pi m} \frac{-\mathrm{i} \boldsymbol{\gamma} \cdot \boldsymbol{r}}{r^{3}} . \tag{44}
\end{equation*}
$$

The above Hamiltonian $\mathcal{H}$ is the last term in the modified Dirac Hamiltonian (right-hand side of equation (3)), approximated for a particle bound in a Coulomb field with the $F_{2}$ form factor evaluated at zero momentum. The electric field $\boldsymbol{E}$ in (44) corresponds to the binding Coulomb interaction. The Hamiltonian (44) describes the modification of the spin-orbit interaction due to the anomalous magnetic moment of the electron.

The nonrelativistic limit of $\mathcal{H}$ is the spin-orbit coupling $H_{\mathrm{fs}}$ given in equation (38), multiplied by a factor $2 F_{2}^{(2)}(0)=\alpha / \pi$ (the additional factor 2 finds an explanation in [30]). The resulting Hamiltonian

$$
\begin{equation*}
H_{\mathrm{eff}}=\frac{\alpha}{\pi} H_{\mathrm{fs}}=\frac{\alpha}{\pi} \frac{Z \alpha}{4 m^{2} r^{3}} \sigma \cdot \boldsymbol{L} \tag{45}
\end{equation*}
$$

takes into account magnetic vertex corrections in the framework of an effective theory. Denoting the variation of the expression (43) mediated by $H_{\text {eff }}$ with the symbol $\delta_{\text {eff }}$, in the spirit of the notation introduced in [13], we obtain the contribution

$$
\begin{equation*}
E_{4 a}=\delta_{\mathrm{eff}}\left\{-\frac{2 \alpha}{3 \pi m} \int_{0}^{\epsilon} \mathrm{d} \omega \omega\langle\phi| p \frac{1}{H-(E-\omega)} p|\phi\rangle\right\} . \tag{46}
\end{equation*}
$$

Following the notation introduced in [11, 12], the contribution $E_{4 a}$ is the sum of the finestructure effects created by the wavefunction-correction $F_{\delta \phi}$, the first relativistic correction to the energy $F_{\delta E}$, and the correction due to the relativistic Hamiltonian $F_{\delta H}$, each multiplied by a factor $\alpha / \pi$. That is to say, the final result for $E_{4}$ is

$$
\begin{equation*}
E_{4 a}=\left(\frac{\alpha}{\pi}\right)^{2}(Z \alpha)^{4} \frac{m}{n^{3}}\left(\Delta_{\mathrm{fs}} F_{\delta \phi}+\Delta_{\mathrm{fs}} F_{\delta E}+\Delta_{\mathrm{fs}} F_{\delta H}\right) . \tag{47}
\end{equation*}
$$

There is a further correction to the nonrelativistic effective coupling to the radiation field due to the 'anomalous spin-orbit Hamiltonian' (44). The correction, in the nonrelativistic limit, can be derived by considering a Foldy-Wouthuysen transformation which by definition diagonalizes the Hamiltonian (44) in spinor space and also leads to a transformation of the relativistic current operator $\alpha^{i}$ according to

$$
\begin{equation*}
\alpha^{i} \rightarrow U \alpha^{i} U^{-1} \quad U=\exp \left(-\mathrm{i} \frac{\beta \mathcal{H}}{2 m}\right) \tag{48}
\end{equation*}
$$

Here, $\beta$ and $\alpha^{i}$ are standard Dirac matrices [17], $i$ is a spatial index, and $\mathcal{H}$ is given in (44). The calculation is carried out along the ideas introduced in [11] and leads to the result

$$
\begin{equation*}
\delta \boldsymbol{j}_{4 b}=\frac{\alpha}{\pi} \frac{Z \alpha}{2 m r^{3}} \boldsymbol{\sigma} \times \boldsymbol{r} \tag{49}
\end{equation*}
$$

as a relativistic correction to the electron current which is simply $\alpha^{i}$ in the relativistic formalism and $p^{i} / m$ in the leading nonrelativistic approximation. Again, following the notation introduced in [11, 12], the resulting additional contribution is

$$
\begin{equation*}
E_{4 b}=\left(\frac{\alpha}{\pi}\right)^{2}(Z \alpha)^{4} \frac{m}{n^{3}} \Delta_{\mathrm{fs}} F_{\delta y} . \tag{50}
\end{equation*}
$$

The sum of (47) and (50) is just the $(Z \alpha)^{6}$-component of the fine-structure difference of the one-loop self-energy from [11, 12], multiplied by an additional factor $\alpha / \pi$. It can also be written as
$E_{4}=E_{4 a}+E_{4 b}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{6} m}{n^{3}}\left[-\frac{n^{2}-1}{3 n^{2}} \ln \frac{2 \epsilon}{(Z \alpha)^{2} m}+\frac{n^{2}-1}{n^{2}} \Delta_{\mathrm{fs}} \ell_{4}(n)\right]$
where $\Delta_{\mathrm{fs}} \ell_{4}(n)$ could be interpreted as a relativistic generalization of a Bethe logarithm, which is $n$-dependent. However, a significant numerical fraction of the $n$-dependence can be eliminated if the factor $\left(n^{2}-1\right) / n^{2}$ is taken out of the final result. The evaluation of $\Delta_{\mathrm{fs}} \ell_{4}(n)$
has recently been performed in [31] with improved numerical methods (see e.g. [32]), and the following results have been obtained:

$$
\begin{align*}
& \Delta_{\mathrm{fs}} \ell_{4}(2)=0.512559768(1) \\
& \Delta_{\mathrm{fs}} \ell_{4}(3)=0.511978815(1) \\
& \Delta_{\mathrm{fs}} \ell_{4}(4)=0.516095539(1)  \tag{52}\\
& \Delta_{\mathrm{fs}} \ell_{4}(5)=0.519976941(1)
\end{align*}
$$

where the uncertainty is due to numerical integration.
There is, as stated above, a further correction due to the explicit modification of the transition current due to the anomalous magnetic moment; it can be obtained through the replacement

$$
\begin{equation*}
\alpha^{i} \rightarrow \alpha^{i}+F_{2}(0) \frac{\mathrm{i} \beta \sigma^{i \nu}}{2 m} q_{\nu} \tag{53}
\end{equation*}
$$

and must be considered in addition to the correction (48). A careful consideration of the nonrelativistic limit of this correction to the current, including retardation effects, leads to the result

$$
\begin{equation*}
\delta \boldsymbol{j}_{5}=\frac{\alpha}{2 \pi} \frac{Z \alpha}{2 m r^{3}} \sigma \times r \tag{54}
\end{equation*}
$$

Consequently, we find that the correction is effectively $F_{2}(0)$ times the retardation corrections to the transition current $F_{\delta y}$ found in $[11,12]$. We obtain

$$
\begin{equation*}
E_{5}=\left(\frac{\alpha}{\pi}\right)^{2}(Z \alpha)^{4} \frac{m}{n^{3}} \frac{\Delta_{\mathrm{fs}} F_{\delta y}}{2} \tag{55}
\end{equation*}
$$

In analogy with $E_{4}$, this correction can favourably be rewritten as

$$
\begin{equation*}
E_{5}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{6} m}{n^{3}}\left[\frac{n^{2}-1}{n^{2}} \Delta_{\mathrm{fs}} \ell_{5}(n)\right] \tag{56}
\end{equation*}
$$

On the basis of $[11,12,31]$, we obtain

$$
\begin{align*}
& \Delta_{\mathrm{fs}} \ell_{5}(2)=-0.173344868(1) \\
& \Delta_{\mathrm{fs}} \ell_{5}(3)=-0.164776514(1) \\
& \Delta_{\mathrm{fs}} \ell_{5}(4)=-0.162263216(1)  \tag{57}\\
& \Delta_{\mathrm{fs}} \ell_{5}(5)=-0.161165602(1) .
\end{align*}
$$

The final result for the low-energy part is

$$
\begin{equation*}
E_{\mathrm{L}}=E_{4}+E_{5} \tag{58}
\end{equation*}
$$

with $E_{4}$ and $E_{5}$ being given in equations (51) and (56), respectively.
We can now understand why it was possible to join the two contributions with 'mixed' and 'low-and-low' energy virtual photons (ii) and (iii), which were discussed in section 1, into a joint 'low-energy part'. The reason is simple: the effective Hamiltonian (45) has no infrared divergence, because it involves the low-energy limit of the magnetic form factor $F_{2}$, which is infrared safe in one-loop order according to equation (9). Because the main contribution to the quantity $F_{2}(0)$ is caused by hard virtual photons, it is also justified to say that the contribution of 'low-and-low' energy virtual photons vanishes at the order of interest for the current calculation (fine-structure difference). In higher-loop order, the further infrared divergence acquired by $F_{2}$ would lead to an infrared divergence in the effective Hamiltonian constructed in analogy with equation (45); this infrared divergence would have to be attributed to a 'mixed' contribution (one photon of high energy and one low-energy photon).

## 5. Results and conclusions

We have obtained analytic results for higher order correction to the two-loop self-energy of $P$ states in hydrogen-like systems. In our calculation, we have analysed the electron form factors through two-loop order in section 2, and we have split the calculation into a highenergy part with two hard virtual photons discussed in section 3, and a low-energy part with at least one soft virtual photon analysed in section 4. The final result for the contribution to the fine-structure energy difference is obtained by adding the high-energy contributions $E_{1}-E_{3}$ given in equations (23), (27), (36), (41), and the low-energy effects $E_{4}$ and $E_{5}$ from equations (51) and (56). The dependence on $\epsilon$ cancels out in the final result which is the sum of the high-energy part $E_{\mathrm{H}}$ given in equation (42) and the low-energy part $E_{\mathrm{L}}$ defined in equation (58). This is also evident when considering explicitly the equations (36) and (51). The final results for the analytic coefficients of order $\alpha^{2}(Z \alpha)^{6}$ read

$$
\begin{equation*}
\Delta_{\mathrm{fs}} B_{61}=-\frac{n^{2}-1}{3 n^{2}} \tag{59}
\end{equation*}
$$

(see also equation (35)) and

$$
\begin{align*}
& \Delta_{\mathrm{fs}} B_{60}=\left(\frac{227}{2880}+\frac{1}{32 n}-\frac{3}{80 n^{2}}\right)+F_{2}^{(4), S}(0)\left(\frac{487}{720}+\frac{5}{8 n}-\frac{23}{20 n^{2}}\right) \\
&+\frac{n^{2}-1}{n^{2}}\left[-\left(F_{1}^{(4), S}(0)+2 \mathcal{F}_{2}^{\prime(4), S}(0)\right) m^{2}-\frac{5}{18}+\Delta_{\mathrm{fs}} \ell_{4}(n)+\Delta_{\mathrm{fs}} \ell_{5}(n)\right] \tag{60}
\end{align*}
$$

where explicit numerical results for $F_{1}^{\prime(4), S}(0), F_{2}^{(4), S}(0)$ and $\mathcal{F}_{2}^{\prime(4), S}(0)$ can be found in equations (17), (18) and (19), respectively. This result refers to the pure self-energy diagrams in figure 1. The result reads numerically for the principal quantum numbers $n=2-5$,

$$
\begin{align*}
& \Delta_{\mathrm{fs}} B_{60}(2)=-0.361196470(1)  \tag{61}\\
& \Delta_{\mathrm{fs}} B_{60}(3)=-0.411156068(1)  \tag{62}\\
& \Delta_{\mathrm{fs}} B_{60}(4)=-0.419926624(1)  \tag{63}\\
& \Delta_{\mathrm{fs}} B_{60}(5)=-0.419832876(1) . \tag{64}
\end{align*}
$$

If it is desired to add in the combined self-energy vacuum-polarization diagram from figure 2 , then the form-factor results from equations (6), (8) and (11) instead of the pure self-energy results given in equations (17)-(19) have to be used in evaluating (60). When including the combined self-energy vacuum-polarization diagram from figure 2 , there is no further lowenergy contribution, so that the alternative set of numerical values for the form factors from equations (6), (8) and (11) fully takes into account the additional effect of the diagram in figure 1 on the fine structure in the order of $\alpha^{2}(Z \alpha)^{6}$.

It is perhaps worth mentioning that for the one-loop self-energy, analytic coefficients are known only up to the order of $\alpha(Z \alpha)^{6}$ [12], but the remaining uncertainty is removed by recent nonperturbative numerical calculations [9, 31, 33]. For the two-loop effect, the ( $Z \alpha$ )expansion converges more rapidly than for the one-loop effect in absolute frequency units because of the additional radiative factor $\alpha / \pi$ which decreases the overall size of the effect.

It is hoped that the analytic calculations for low nuclear charge number $Z$ will be supplemented in the future by an accurate numerical treatment of the two-loop self-energy problem (see also related recent work in the high- $Z$ region [34-36]). This presupposes that the considerable numerical problems in the domain of small nuclear charge could be solved
by adequate numerical methods, and that the further problem of the increased computational demand of the two-loop effect in comparison to the one-loop problem [9,33] can be tackledpossibly by massively parallel computer architectures. Note, however, that the most accurate theoretical predictions could only be reached by combining the numerical and analytic results. The reason is the following: all numerical calculations are performed in the non-recoil limit which is the limit of infinite nuclear mass. This is not quite sufficient for an accurate theoretical treatment because the self-energy effect for a bound state depends genuinely on the ratio of the orbiting particle to the nuclear mass-an effect beyond the recoil correction. For example, the argument of the logarithms in (1) should be replaced according to $\ln \left[(Z \alpha)^{-2}\right] \rightarrow \ln \left[\sigma(Z \alpha)^{-2}\right]$, where $\sigma=m / m_{\mathrm{r}}$ and $m_{\mathrm{r}}$ is the reduced mass [3]. The possibility of including these tiny, but important effects depends crucially on a reliable knowledge of the analytic coefficients in combination with an accurate numerical treatment of the problem.

The analytic results can be used to obtain improved theoretical predictions for the hydrogenic fine structure compared to the previous order- $\alpha^{7}$ calculations [11, 12], because they remove the principal theoretical uncertainty in the order of $\alpha^{8}$ due to the problematic two-loop self-energy which is represented diagrammatically in figure 1. A compilation of the other corrections relevant at the order of $\alpha^{8}$, including but not limited to the vacuum polarization effects, whose evaluation is rather straightforward, will be presented elsewhere. Our calculation illustrates the usefulness of the simplified effective treatments of two-loop effects in the analytic approach based on the modified Dirac Hamiltonian (3) and the ' $\epsilon$ method' (see $[8,11,12,19]$ and appendix A). This aspect highlights, as we believe, the need for systematic, simplified treatments of higher order radiative corrections in bound systems.

In this paper, we primarily address spin-dependent effects in one-electron (hydrogen-like) systems. However, the same effects also contribute to the fine-structure splitting in two-electron (heliumlike) systems. There is currently remarkable interest in improved measurements of the fine-structure splitting in helium and heliumlike atomic systems with low nuclear charge [37]. The effects addressed in this paper contribute to the fine-structure splitting in helium at the level of 100 Hz , which is not much smaller than the current experimental accuracy of about 1 kHz , and allows for an estimate of uncalculated yet higher order contributions.

The results of this paper are a step in a systematic study of the higher order binding corrections to the two-loop Lamb shift of S and P states. The scheme of calculation permits not only a simplified treatment of the problem via a separation into appropriate energy regions for the two virtual photons, but also a clear identification of the spin-independent and spindependent contributions to the self-energy. The results will therefore be directly applicable to the total coefficients for P states once the spin-independent parts (the two-loop Bethe logarithms) are calculated. These are currently under study.

## Acknowledgments

UDJ acknowledges helpful conversations with Professor Gerhard Soff. This work was supported in part by a research grant from the Polish Committee for Scientific Research under contract no 2P03B 05718.

## Appendix A. The ' $\epsilon$ method'

We discuss here, by way of example, the $\epsilon$ method employed in the analytic calculation of self-energy effects in bound systems. This method is very suitable [11, 12] for the separation
of the two different energy scales for virtual photons: the nonrelativistic domain, in which the virtual photon assumes values of the order of the atomic binding energy, and the relativistic domain, in which the virtual photon assumes values of the order of the electron rest mass. Different approximation schemes and different asymptotic expansions are adequate for the two different domains. Without these approximations and expansions, the analytic evaluation of either the high- or low-energy part would not be feasible. At the same time, the model example discussed in this appendix is meant to illustrate the usefulness of the ' $\epsilon$ method' in a more general context.

We will consider here a model problem with only one 'virtual photon'. The separation into high- and low-energy photons necessitates the temporary introduction of a parameter $\epsilon$; the dependence on $\epsilon$ cancels when the high- and low-energy parts are added together. We have

$$
\begin{align*}
& \text { nonrelativistic domain } \ll \epsilon \ll \text { electron rest mass }  \tag{A.1}\\
& (Z \alpha)^{2} m_{\mathrm{e}} \ll \epsilon<m_{\mathrm{e}} \tag{A.2}
\end{align*}
$$

where $\alpha$ is the fine structure constant and $Z$ is the nuclear charge. The high-energy part is associated with photon energies $\omega>\epsilon$ and low-energy part is associated with photon energies $\omega<\epsilon$.

In order to illustrate the procedure, we discuss a simple, one-dimensional example: the evaluation of

$$
\begin{equation*}
I(\beta)=\int_{0}^{1} \sqrt{\frac{\omega^{2}+\beta^{2}}{1-\omega^{2}}} \mathrm{~d} \omega . \tag{A.3}
\end{equation*}
$$

where the integration variable $\omega$ might be interpreted as the 'energy' of a 'virtual photon'. The integral $I$ can be expressed in terms of special functions,

$$
\begin{equation*}
I(\beta)=\beta E\left(-\frac{1}{\beta^{2}}\right)=\beta \frac{\pi}{2}{ }_{2} F_{1}\left(-\frac{1}{2}, \frac{1}{2} ; 1 ;-\frac{1}{\beta^{2}}\right) \tag{A.4}
\end{equation*}
$$

where $E$ is the complete elliptic integral of the second kind, and ${ }_{2} F_{1}$ denotes a hypergeometric function. An alternative integral representation reads $I(\beta)=\int_{0}^{\pi / 2} \sqrt{\beta^{2}+\sin ^{2}(\omega)} \mathrm{d} \omega$.

The purpose of the calculation is to derive a semi-analytic expansion of $I(\beta)$ in powers of $\beta$ and $\ln \beta$. The fine structure constant $\alpha$ takes the role of the expansion parameter $\beta$ in actual self-energy calculations. We discuss first the 'high-energy part' of the calculation. It is given by the expression

$$
\begin{equation*}
I_{\mathrm{H}}(\beta)=\int_{\epsilon}^{1} \sqrt{\frac{\omega^{2}+\beta^{2}}{1-\omega^{2}}} \mathrm{~d} \omega . \tag{A.5}
\end{equation*}
$$

For $\omega>\epsilon$, we may expand

$$
\begin{equation*}
\sqrt{\omega^{2}+\beta^{2}}=\omega+\frac{\beta^{2}}{2 \omega}+\frac{\beta^{4}}{8 \omega^{3}}+\mathcal{O}\left(\beta^{6}\right) \tag{A.6}
\end{equation*}
$$

but this expansion is not applicable in higher orders to the domain $0<\omega<\epsilon$ because of the appearance of inverse powers of $\omega$ (analogous to an 'infrared divergence' in QED).

The separation parameter $\epsilon$ acts an infrared regulator. After expanding in $\beta$ (see equation (A.6)), the resulting integrals in each order of $\beta$ can be evaluated analytically. Subsequently, we expand every term in the $\beta$-expansion in powers of $\epsilon$ up to the order $\epsilon^{0}$, i.e. we keep only the divergent and constant terms in $\epsilon$. The result is

$$
\begin{align*}
I_{\mathrm{H}}(\beta, \epsilon)=1+ & \beta^{2}\left\{\frac{1}{2} \ln \left(\frac{2}{\epsilon}\right)+\mathcal{O}(\epsilon)\right\}+\beta^{4}\left\{-\frac{1}{16 \epsilon^{2}}-\frac{1}{16} \ln \left(\frac{2}{\epsilon}\right)+\frac{1}{32}+\mathcal{O}(\epsilon)\right\} \\
& +\beta^{6}\left\{\frac{1}{64 \epsilon^{4}}+\frac{1}{64 \epsilon^{2}}+\frac{3}{128} \ln \left(\frac{2}{\epsilon}\right)-\frac{7}{512}+\mathcal{O}(\epsilon)\right\}+\mathcal{O}\left(\beta^{8}\right) \tag{A.7}
\end{align*}
$$

Here, the ' $\mathcal{O}$ '-symbol identifies a contribution for which $\mathcal{O}(x) / x \rightarrow$ const as $x \rightarrow 0$, whereas the ' o '-symbol identifies the weaker requirement $\mathrm{o}(x) \rightarrow 0$ as $x \rightarrow 0$; this is consistent with the standard notation (see e.g. [38]).

The contribution $I_{\mathrm{H}}(\beta)$ corresponds to the 'high-energy part' in analytic self-energy calculations, where the propagator of the bound electron is explicitly expanded in powers of the fine structure constant $\alpha$. Now we turn to the 'low-energy part'. The expression for the low-energy part $(0<\omega<\epsilon)$ reads

$$
\begin{equation*}
I_{\mathrm{L}}(\beta)=\int_{0}^{\epsilon} \sqrt{\frac{\omega^{2}+\beta^{2}}{1-\omega^{2}}} \mathrm{~d} \omega \tag{A.8}
\end{equation*}
$$

The expansion (A.6) is not applicable in this energy domain; we therefore have to keep the numerator of the integrand $\sqrt{\omega^{2}+\beta^{2}}$ in unexpanded form. However, we can expand the denominator $\sqrt{1-\omega^{2}}$ of the integrand in powers of $\omega$; because $0<\omega<\epsilon$ (with $\epsilon$ small), this expansion in $\omega$ is in fact an expansion in $\beta$-although the situation is somewhat problematic in the sense that every term in the $\omega$-expansion gives rise to terms of arbitrarily high order in the $\beta$-expansion (see also equation (A.10) below).

The term $\sqrt{\omega^{2}+\beta^{2}}$ is analogous to the Schrödinger-Coulomb propagator in the selfenergy calculation which has to be kept in unexpanded form, whereas the expansion

$$
\begin{equation*}
\frac{1}{\sqrt{1-\omega^{2}}}=1+\frac{\omega^{2}}{2}+\frac{3}{8} \omega^{4}+\mathcal{O}\left(\omega^{6}\right) \tag{A.9}
\end{equation*}
$$

corresponds to the expansion into the $(Z \alpha)$-expansion in the low-energy part.
Every term in the expansion (A.9) gives rise to arbitrarily high order corrections in $\beta$, but it starts with the power $\omega^{n} \rightarrow \beta^{n+2}$. For example, we have for the leading term of order $\omega^{0}=1$ from equation (A.9)

$$
\begin{align*}
& \int_{0}^{\epsilon} \sqrt{\omega^{2}+\beta^{2}} \mathrm{~d} \omega=\beta^{2}\left\{\frac{1}{2} \ln \left(\frac{2}{\beta} \epsilon\right)+\frac{1}{4}+\mathcal{O}(\epsilon)\right\}+\beta^{4}\left\{\frac{1}{16 \epsilon^{2}}+\mathcal{O}(\epsilon)\right\} \\
&+\beta^{6}\left\{-\frac{1}{64 \epsilon^{4}}+\mathcal{O}(\epsilon)\right\}+\mathcal{O}\left(\beta^{8}\right) \tag{A.10}
\end{align*}
$$

Note that the terms generated in the orders $\beta^{4}$ and $\beta^{6}$ are needed to cancel divergent contributions in respective orders of $\beta$ from the high-energy part given in equation (A.6). The term of order $\omega^{2}$ from (A.9) results in

$$
\begin{equation*}
\frac{1}{2} \int_{0}^{\epsilon} \omega^{2} \sqrt{\omega^{2}+\beta^{2}} \mathrm{~d} \omega=\beta^{4}\left\{-\frac{1}{16} \ln \left(\frac{2}{\beta} \epsilon\right)+\frac{1}{64}+\mathcal{O}(\epsilon)\right\}+\beta^{6}\left\{-\frac{1}{64 \epsilon^{2}}+\mathcal{O}(\epsilon)\right\}+\mathcal{O}\left(\beta^{8}\right) \tag{A.11}
\end{equation*}
$$

Altogether, we obtain for the low-energy part,

$$
\begin{align*}
& I_{\mathrm{L}}(\beta, \epsilon)=\beta^{2}\left\{\frac{1}{2} \ln \left(\frac{2}{\beta} \epsilon\right)+\frac{1}{4}+\mathcal{O}(\epsilon)\right\}+\beta^{4}\left\{\frac{1}{16 \epsilon^{2}}-\frac{1}{16} \ln \left(\frac{2}{\beta} \epsilon\right)+\frac{1}{64}+\mathcal{O}(\epsilon)\right\} \\
&+\beta^{6}\left\{-\frac{1}{64 \epsilon^{4}}-\frac{1}{64 \epsilon^{2}}+\frac{3}{128} \ln \left(\frac{2}{\beta} \epsilon\right)-\frac{5}{512}+\mathcal{O}(\epsilon)\right\}+\mathcal{O}\left(\beta^{8} \ln \beta\right) \tag{A.12}
\end{align*}
$$

When the high-energy (A.7) and low-energy parts (A.12) are added, the dependence on $\epsilon$ cancels, and we have

$$
\begin{gather*}
I(\beta)=I_{\mathrm{H}}(\beta, \epsilon)+I_{\mathrm{L}}(\beta, \epsilon)=1+\beta^{2}\left\{\frac{1}{2} \ln \left(\frac{4}{\beta}\right)+\frac{1}{4}\right\}+\beta^{4}\left\{-\frac{1}{16} \ln \left(\frac{4}{\beta}\right)+\frac{3}{64}\right\} \\
+\beta^{6}\left\{\frac{3}{128} \ln \left(\frac{4}{\beta}\right)-\frac{3}{128}\right\}+\mathcal{O}\left(\beta^{8} \ln \beta\right) \tag{A.13}
\end{gather*}
$$

In order to illustrate the analogy with the self-energy calculation presented here, we would like to point out that the dependence on $\epsilon$ cancels out in the final result which is the sum of the high-energy part $E_{\mathrm{H}}$ given in equation (42) and the low-energy part $E_{\mathrm{L}}$ in equation (58).

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[^0]:    ${ }^{4}$ Certain commercial equipment, instruments or materials are identified in this paper to foster understanding. Such identification does not imply recommendation or endorsement by the National Institute of Standards and Technology, nor does it imply that the materials or equipment identified are necessarily the best available for the purpose.

